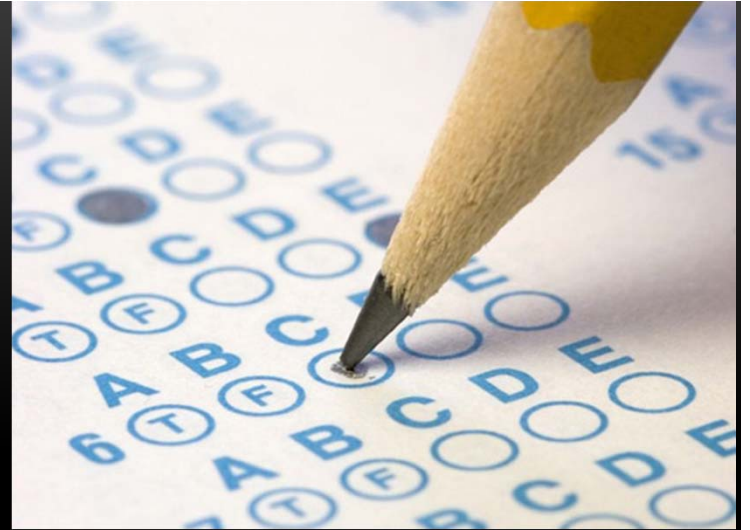


# TEST IS COMING

- Test on Friday, 9/22
  - Ch.1 through Ch.5
  - Multiple choice, bring #2 pencil, eraser, & your ACT testing skills
  - No need for calculator
  - No phones
- Work through the practice test posted on the class webpage, end of chapter stuff in Mastering



# DOPPLER SHIFT APPLIED

## Laboratory spectrum

*Lines at rest wavelengths.*



## Object 1 *Lines redshifted:*

*Object moving away from us.*



## Object 2 *Greater redshift:*

*Object moving away faster than object 1.*



## Object 3 *Lines blueshifted:*

*Object moving toward us.*



## Object 4 *Greater blueshift:*

*Object moving toward us faster than object 3.*

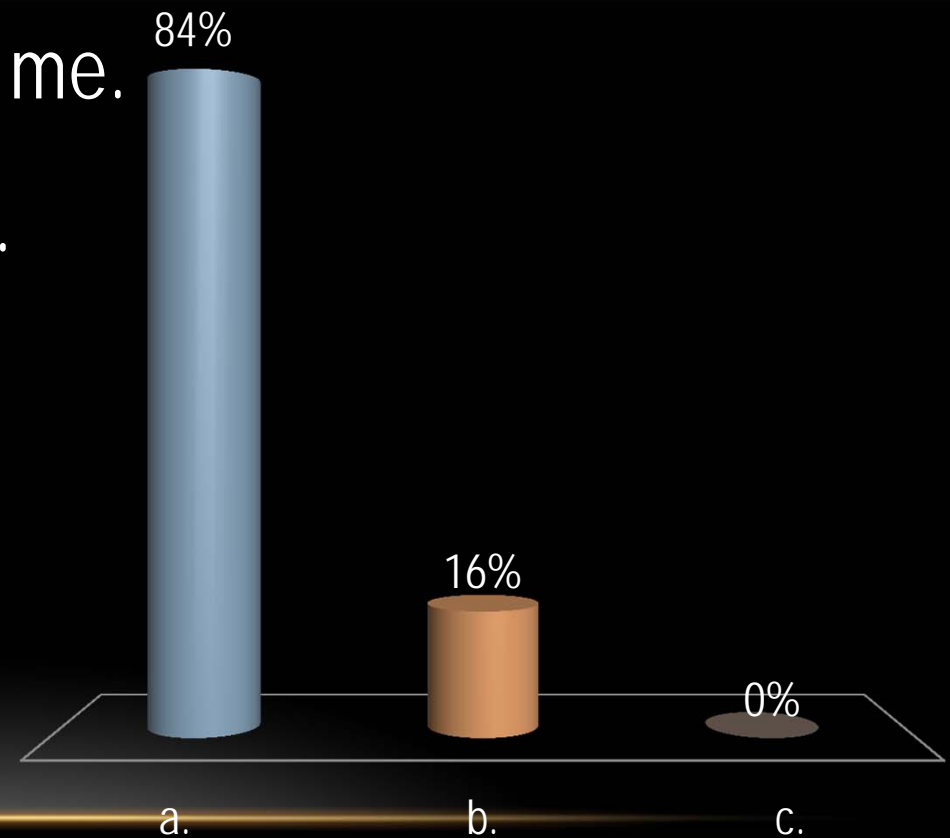


Fig.5.14

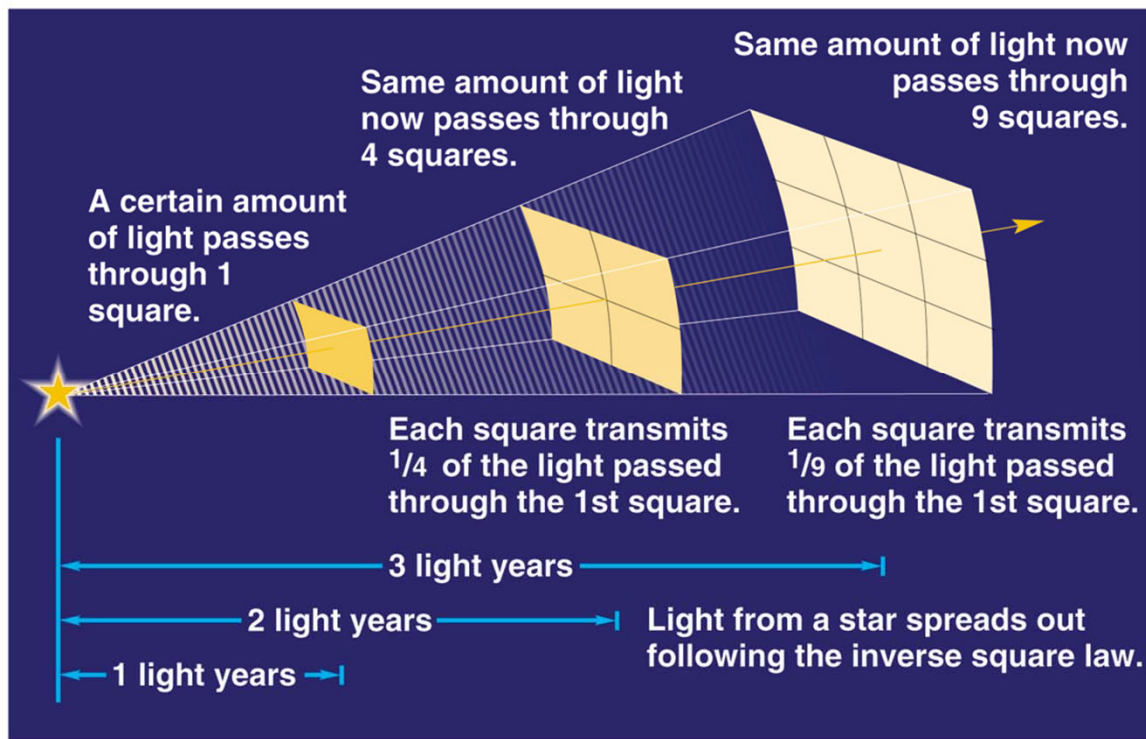
PLAY

I MEASURE A LINE IN THE LAB AT 500.7 NM. THE SAME LINE IN A STAR HAS WAVELENGTH 502.8 NM. WHAT CAN I SAY ABOUT THIS STAR?

- ✓ a. It is moving away from me.
- b. It is moving toward me.
- c. It has unusually long spectral lines.




# INVERSE SQUARE LAW



- How bright something is can also be measured
- Intensity, like gravity, falls off as  $1/r^2$  where  $r$  is distance
- So, further things are fainter than they would be if closer



# LIGHT & SPECTRA

- Light (and other EM waves) is the only thing we have to work with in Astronomy
  - We can measure:
    - Direction
    - Wavelength (spectra)
    - Intensity
- 

# DIRECTION

- Where does the light appear to be coming from?
  - This star, some other star, our galaxy, another galaxy?


# WAVELENGTH

- A spectrum is what you get when you take all the light coming from something and split it up by wavelength (color)  $\lambda$
- Used to tell:
  - Chemical composition
  - Temperature
  - Velocity towards and away from us

# INTENSITY

- How bright something is
- Follows inverse square law
- If we know how bright something is at a known distance, and we see something similar at an unknown distance, how bright it is will tell us how far away it is
  - "Standard Candle"

# LOTS OF INFORMATION!

- EM waves are packed with the information which allow us to decode what is happening out there
  - We will use these tools for the rest of the semester to understand our solar system, galaxy, and universe
- 

# PRACTICE QUIZ

- Three “Main Sequence” stars are:
  - Proxima Centauri, which is red
  - Sirius, which is white
  - Our Sun, which is yellow
- Write down the stars in order from coolest to hottest
- Which one puts out the least energy?
- Which one emits the most Ultraviolet light?

Without clickers on this one: write answers in your notes instead.  
(couldn't figure out how to get a more complicated  
series of questions like this to work with clickers)

# ANSWERS:

- Proxima Centauri, the Sun, Sirius
- Proxima Centauri puts out the least amount of energy
- Sirius emits the most UV
- Note that Betelgeuse, a Red Supergiant, puts out more energy than any of these – why?
  - Size matters! The temperature determines how much energy per square meter, but Betelgeuse is as big across as the orbit of Mars, so it has a Lot of Square Meters



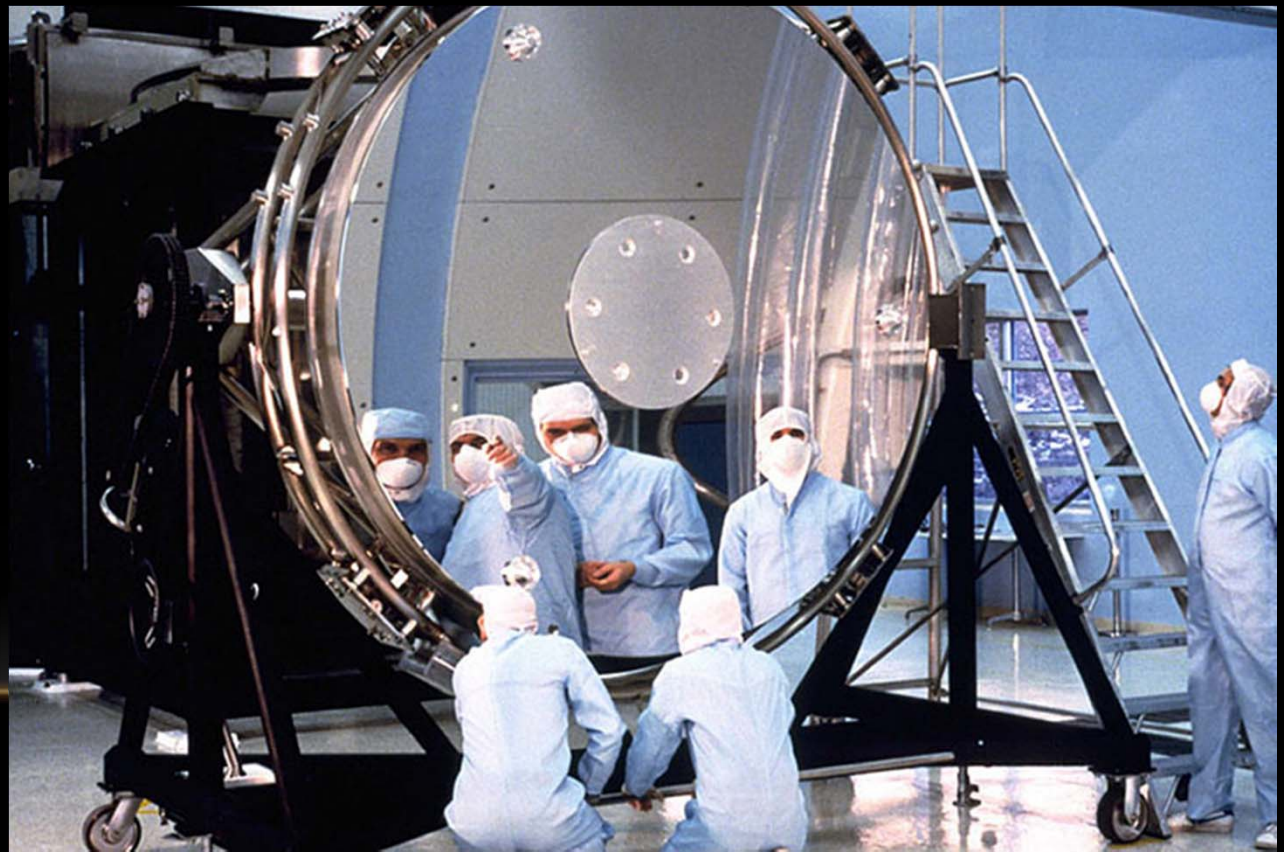
# TELESCOPES

- Let us see fainter things by collecting more light
- Let us see smaller details by improving angular resolution
- Let us see all the electromagnetic waves which aren't visible light

# LIGHT BUCKETS

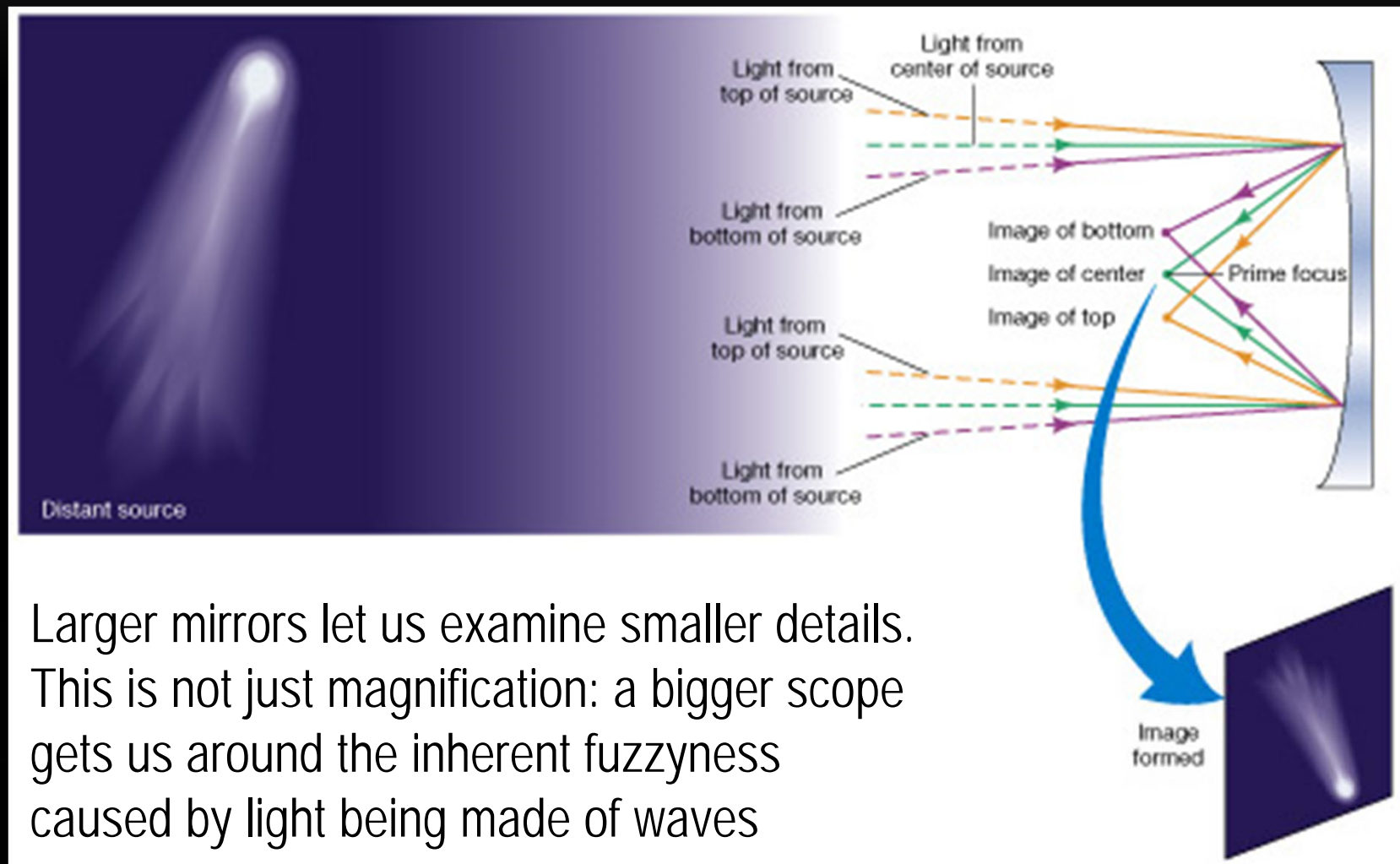
- Main job: collect more light by simply being bigger than our own pupils
- See dimmer things

Workers inspect the Hubble's 8ft (2.4m) mirror before it was launched



PLAY

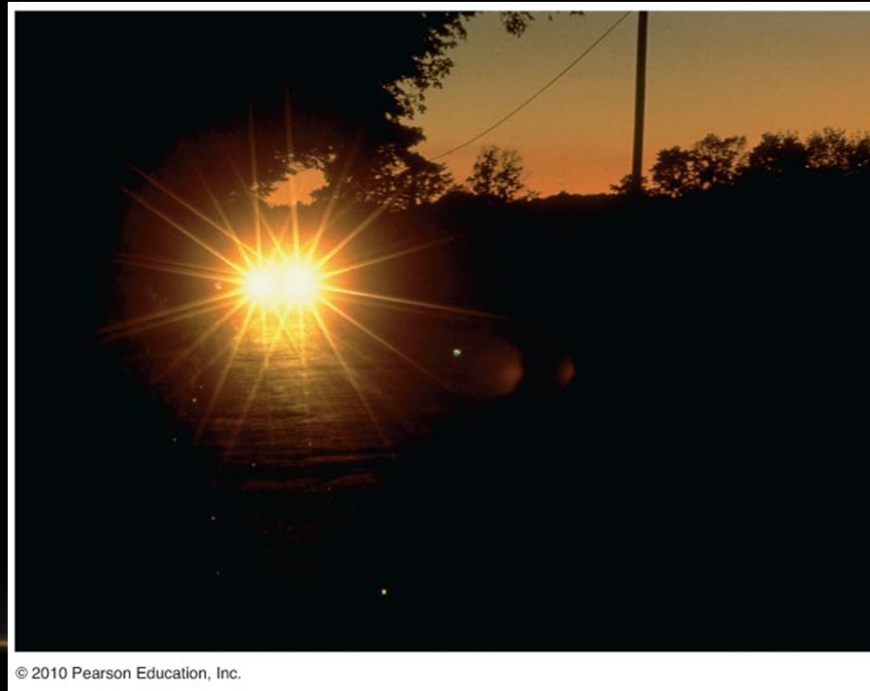
# IMAGE FORMATION



Larger mirrors let us examine smaller details.  
This is not just magnification: a bigger scope  
gets us around the inherent fuzziness  
caused by light being made of waves

# ANGULAR RESOLUTION

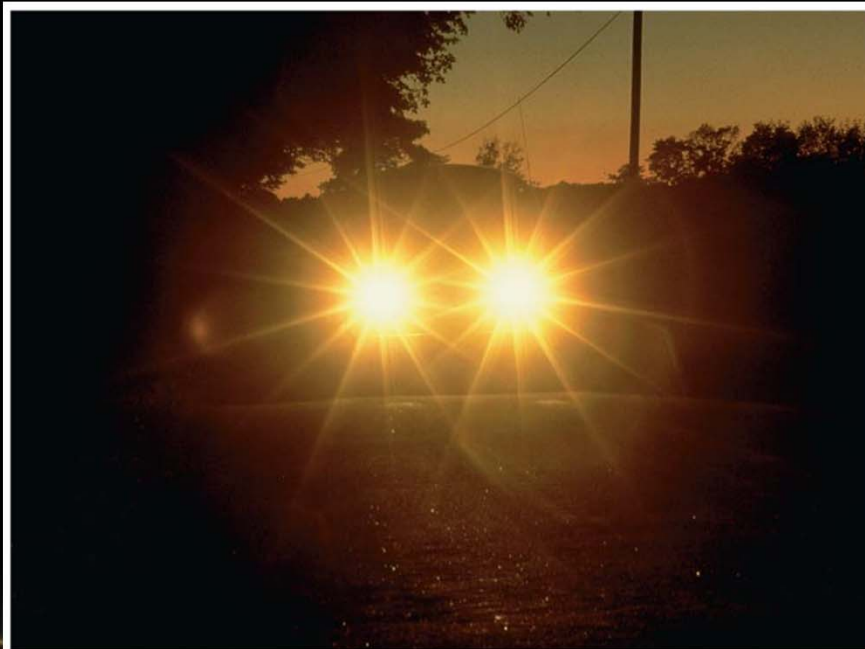
- How close together (in angle) can things get before you can't tell them apart?



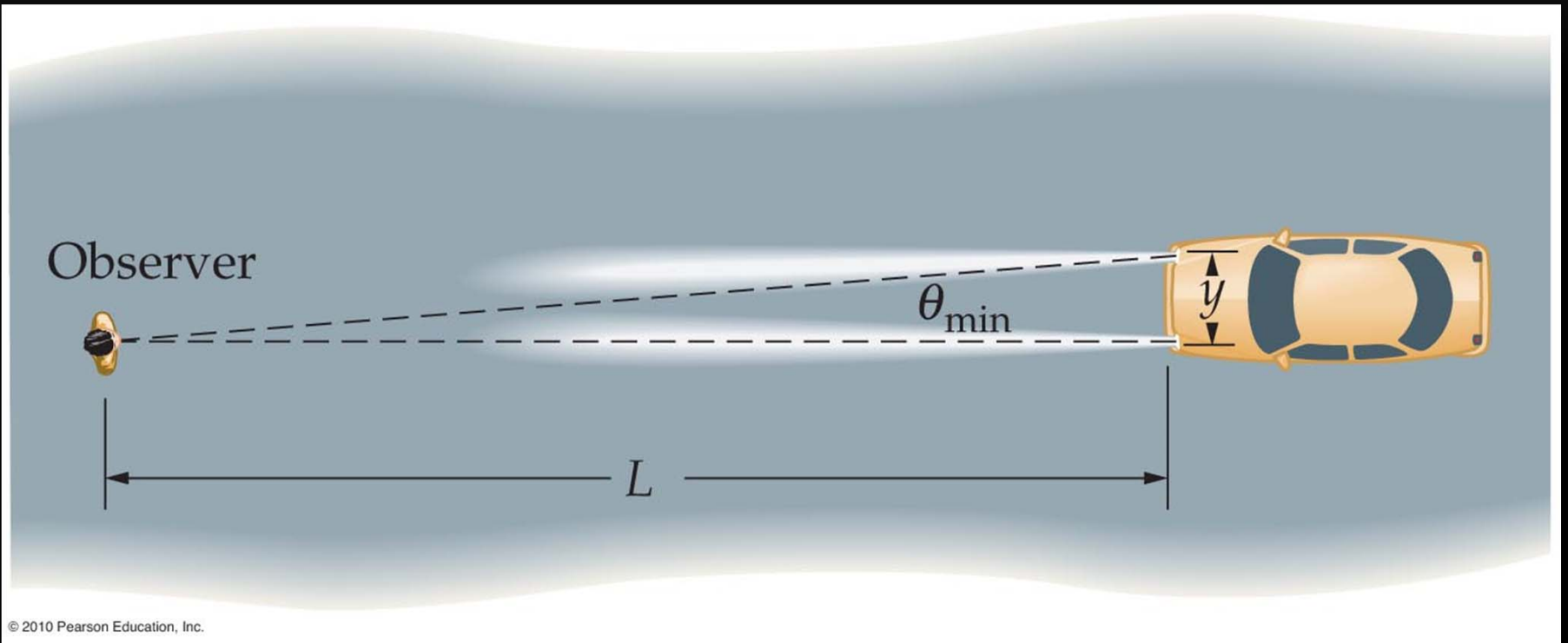
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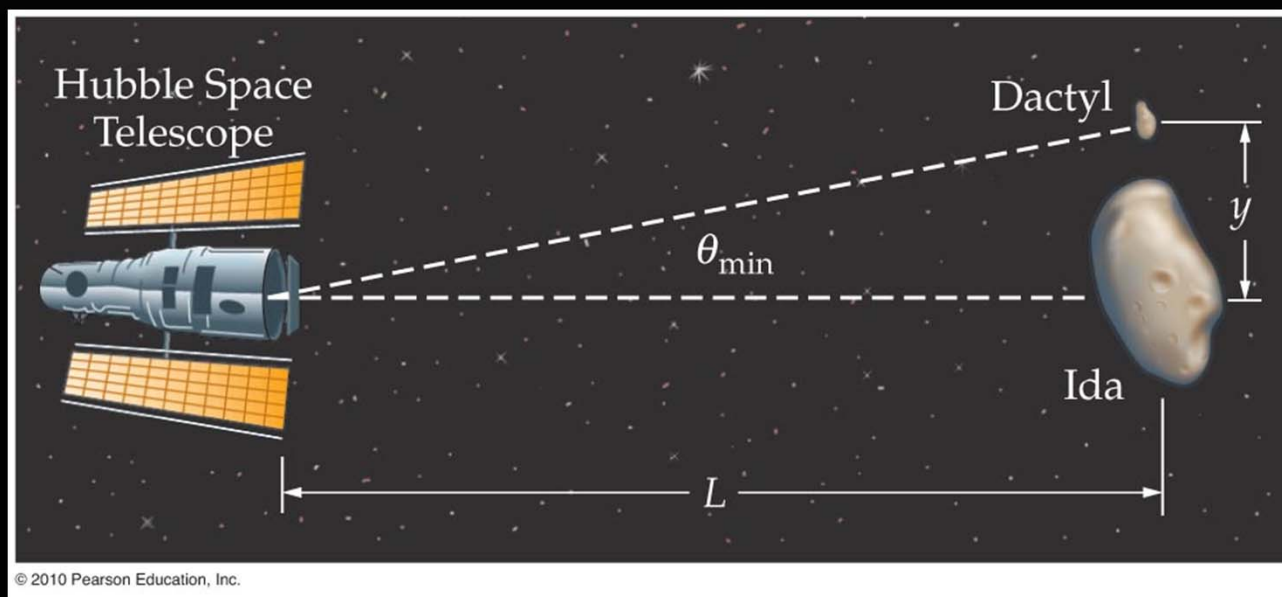


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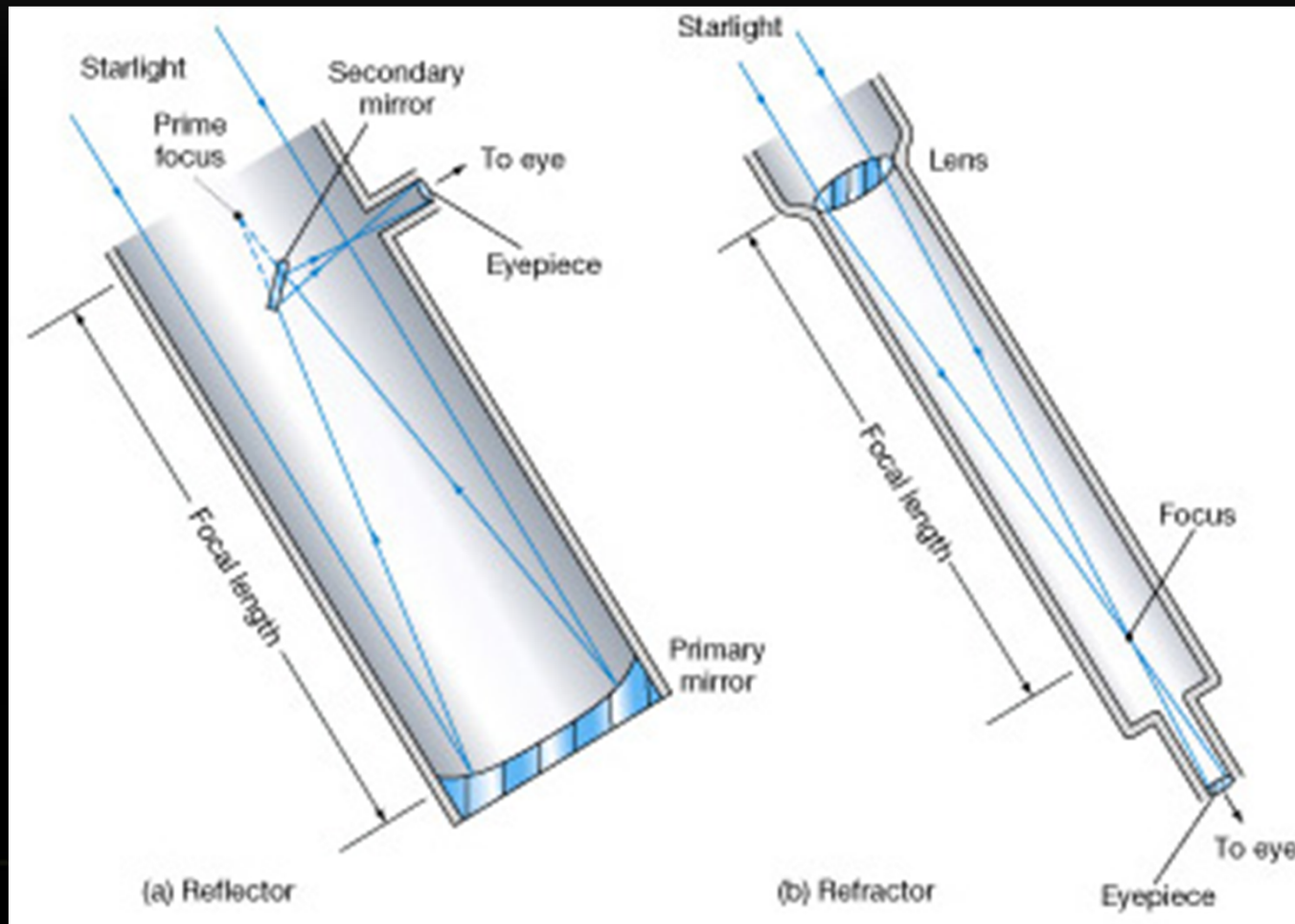
PLAY





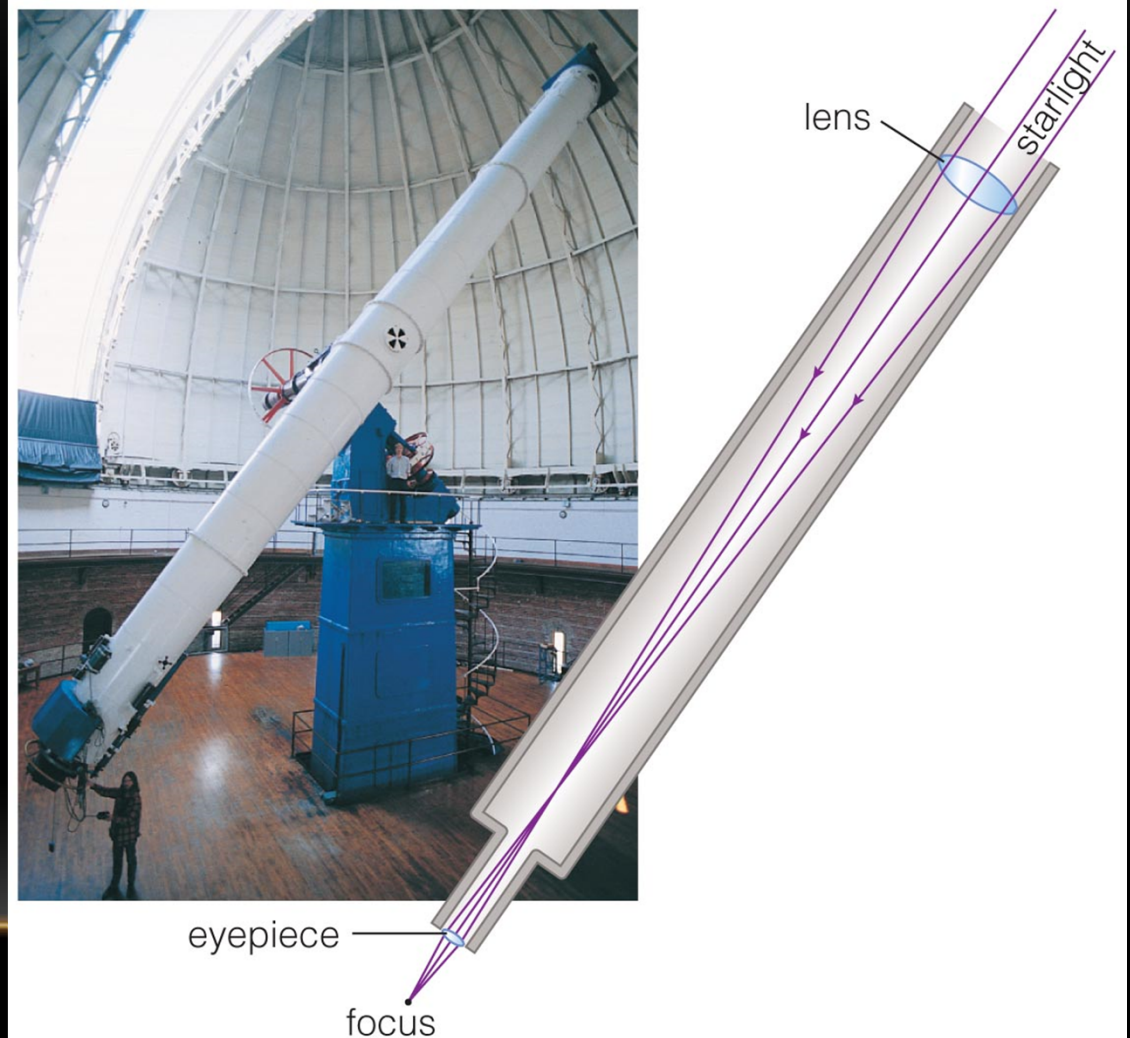
PLAY

# REFRACTOR VS. REFLECTOR



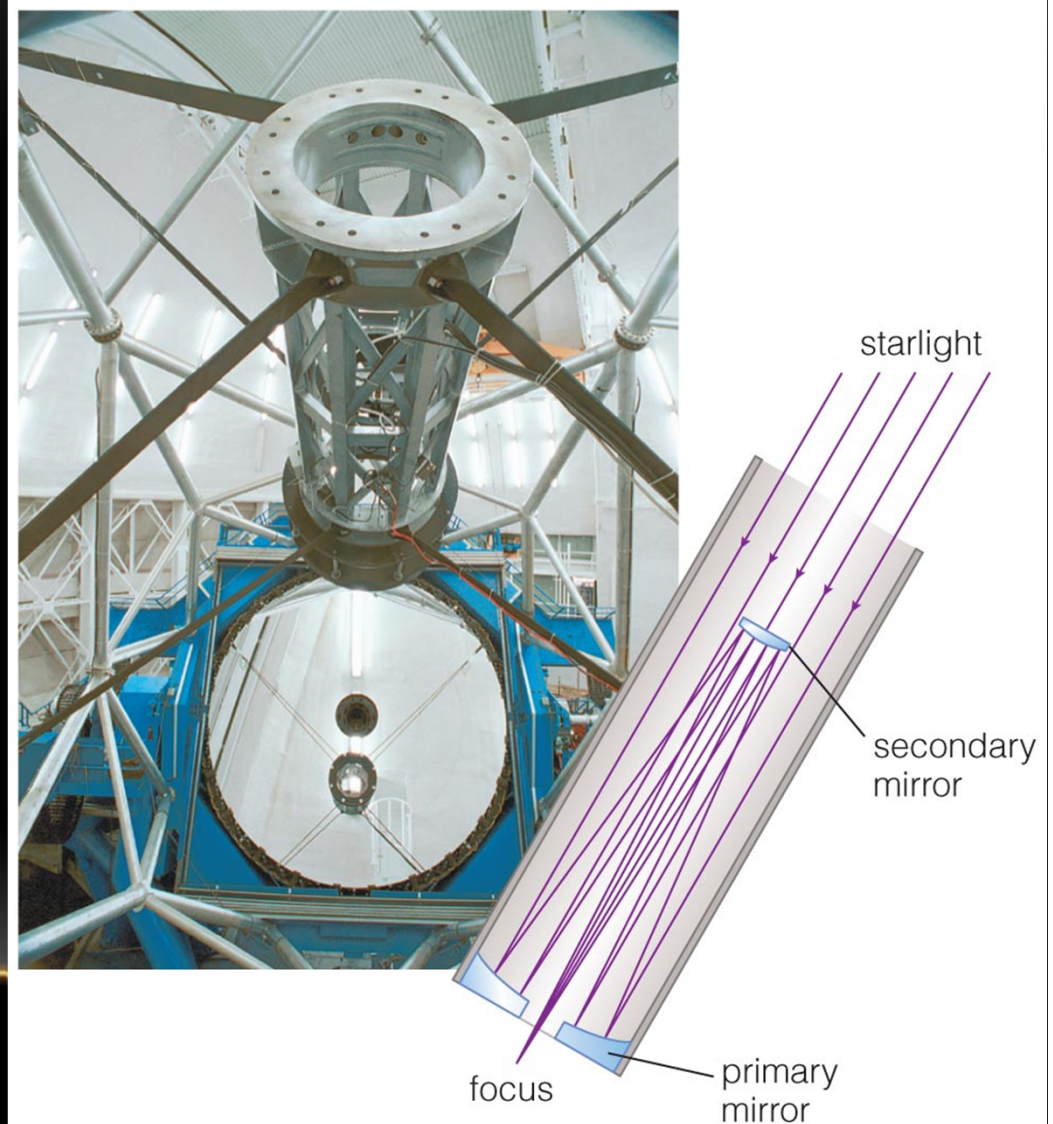
# REFRACTORS

- Use lenses
- Mechanically hard to make these very large: this 1-m in Yerkes, WI is the largest



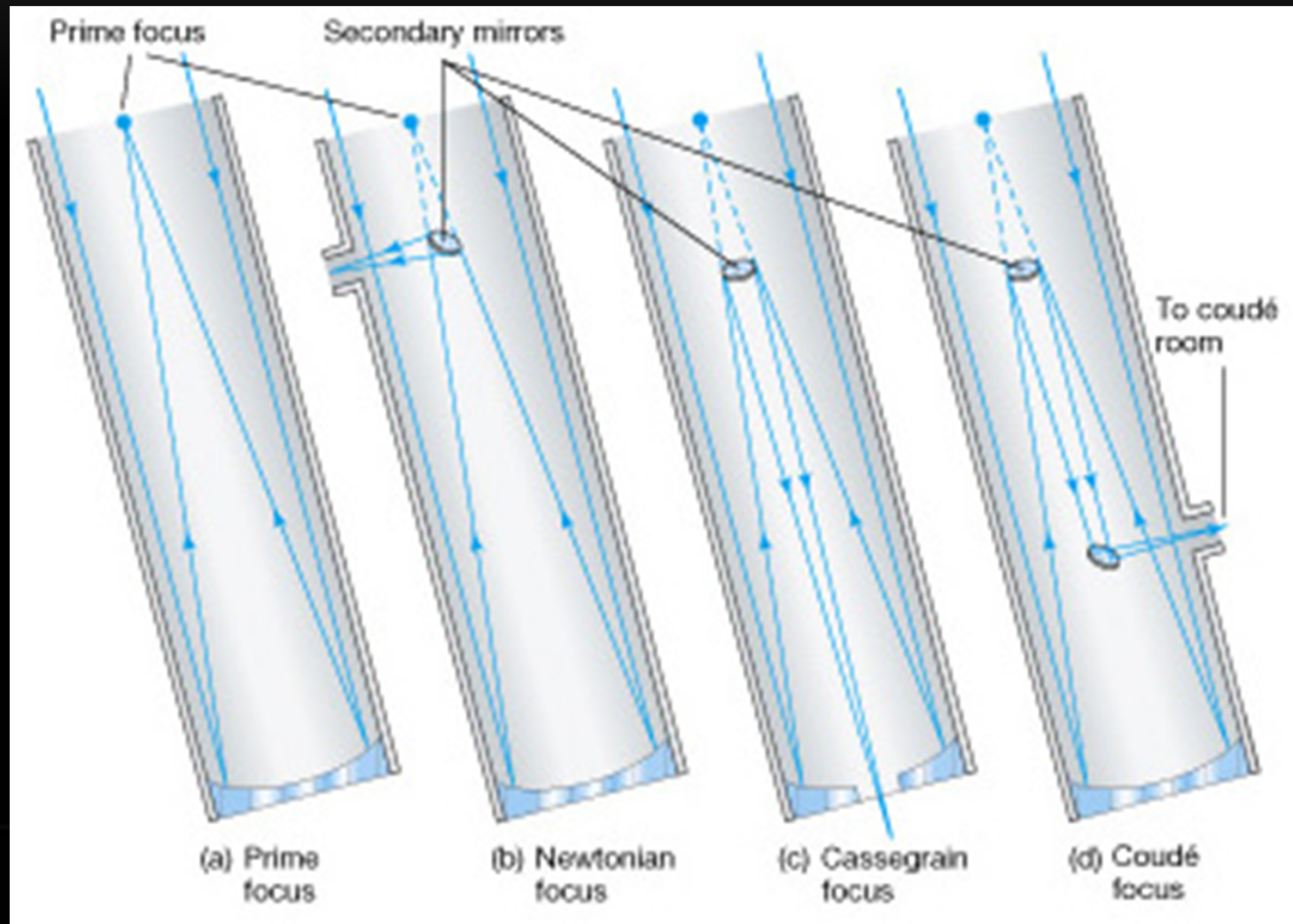
# REFLECTORS

- Use mirrors
- Much easier to build huge mirrors
- All research-grade scopes are reflectors
  - This one is the 8m Gemini North

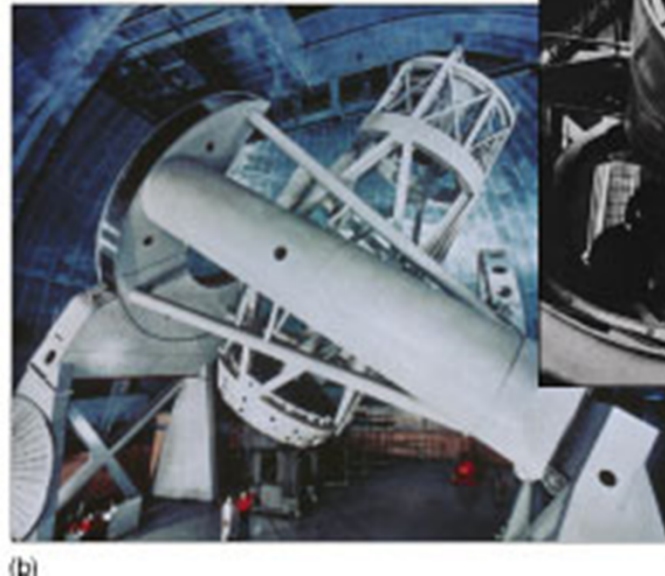
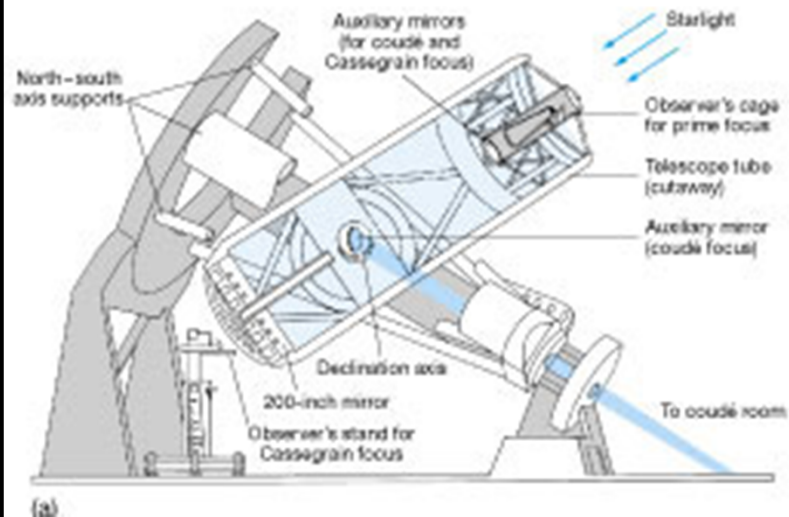




# REFLECTORS



# PALOMAR 200 INCH



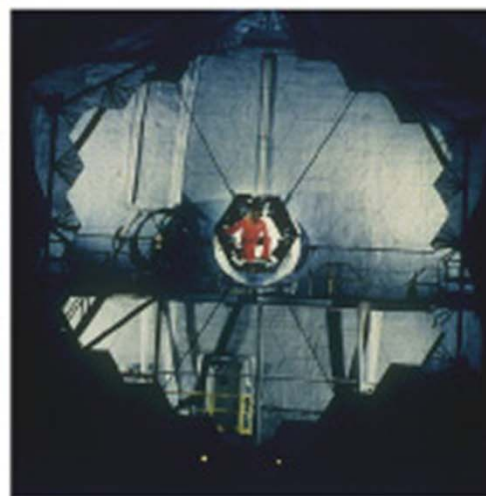
# KECK



(a)



Robert Van Geen



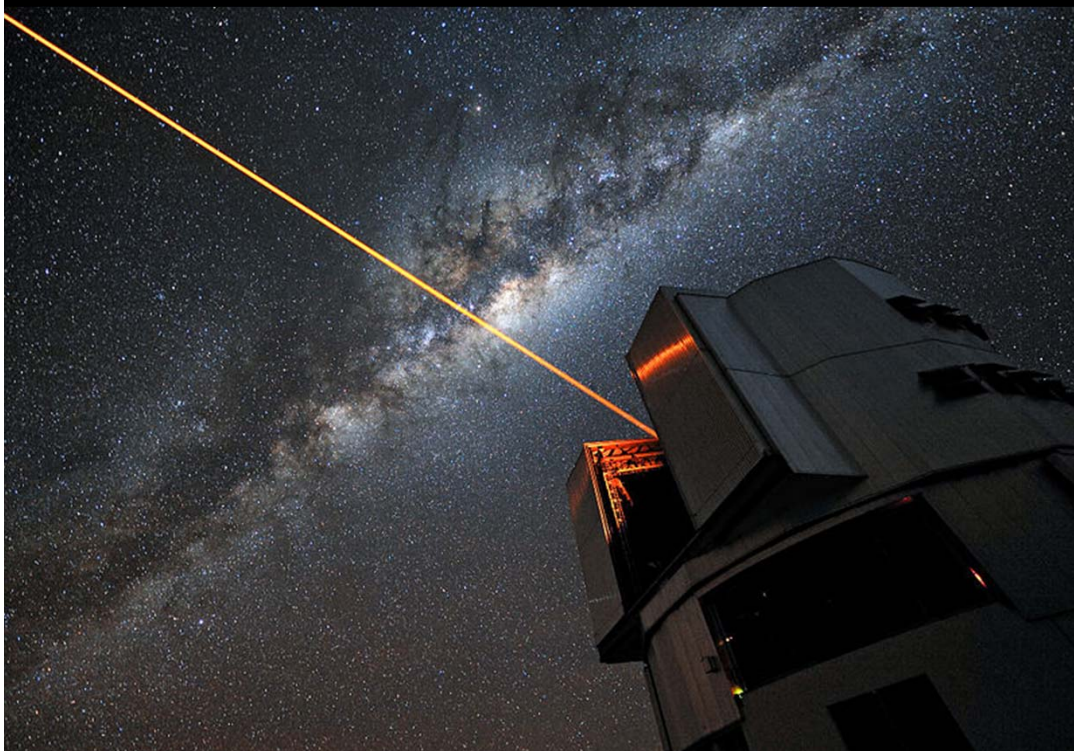
(b)





# LASERS AND ADAPTIVE OPTICS

- Scopes like Keck can use an artificial star to see what the air is doing, then warp their mirrors to “de-twinkle” their images



The laser excites sodium atoms high in the atmosphere, causing them to emit light

ESO's VLT (Chile)

# BIGGER IS BETTER

More light

Better  
Resolution



(a)



(b)



(a)



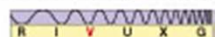
(b)



(c)

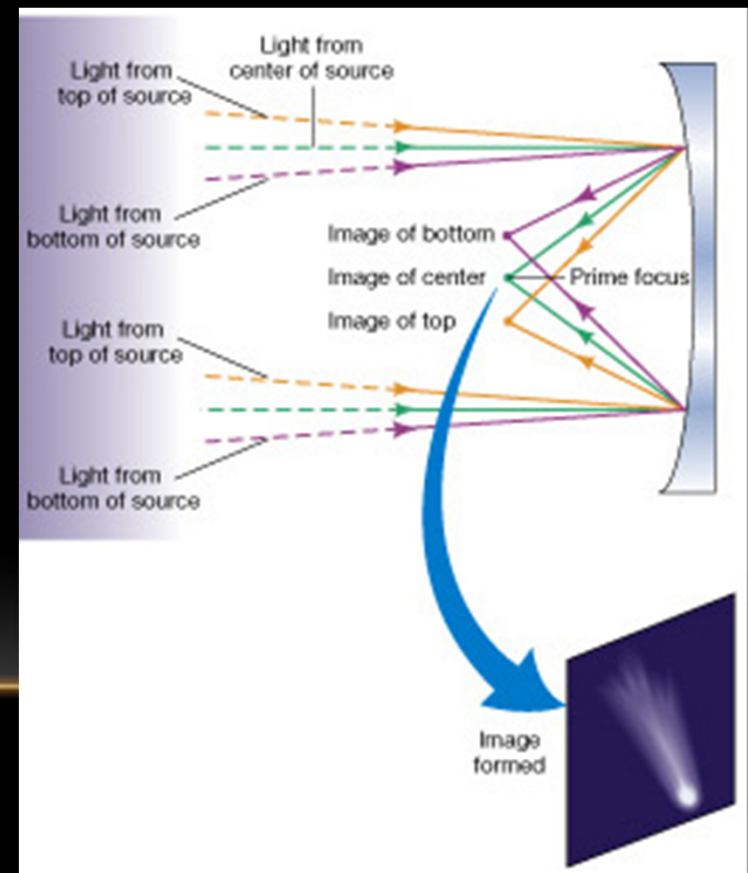


(d)



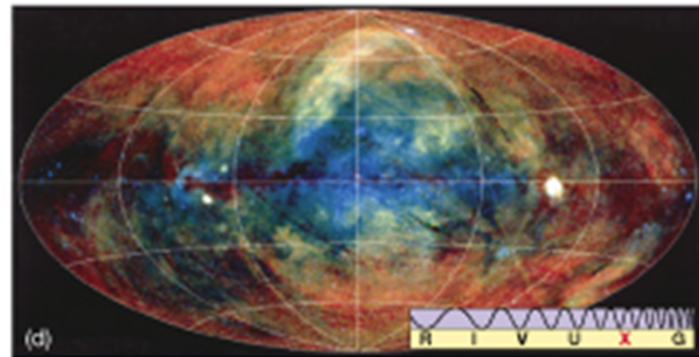
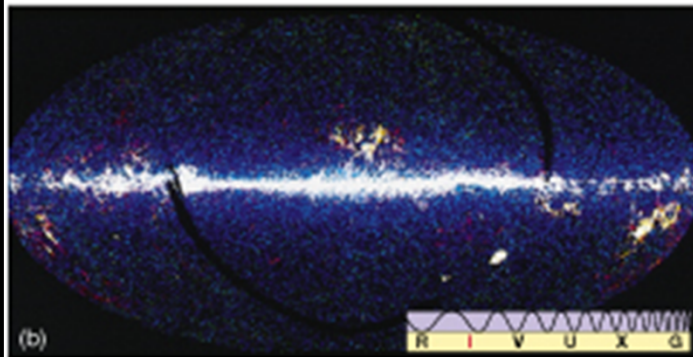
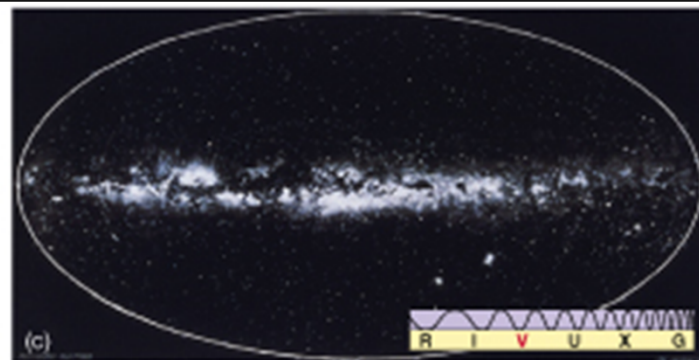
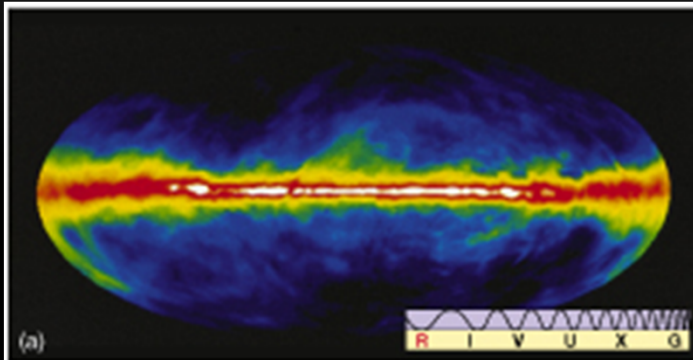
# MAGNIFICATION?

- Not so important. Change the last little eyepiece lens. That last lens just is inspecting the image produced by the main lens or mirror, like a magnifying glass

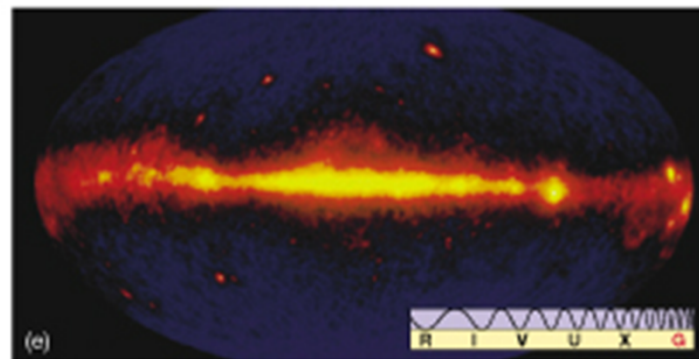




# OTHER E-M WAVES



The Milky Way  
seen in many  
wavelengths



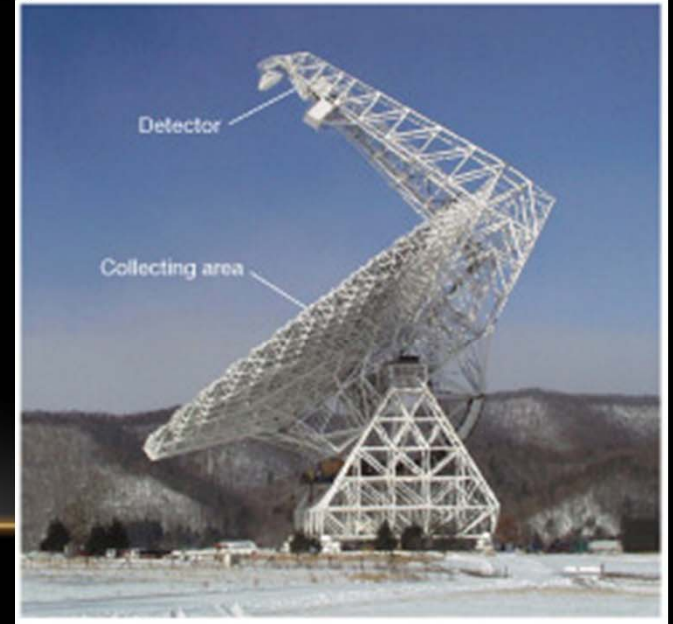
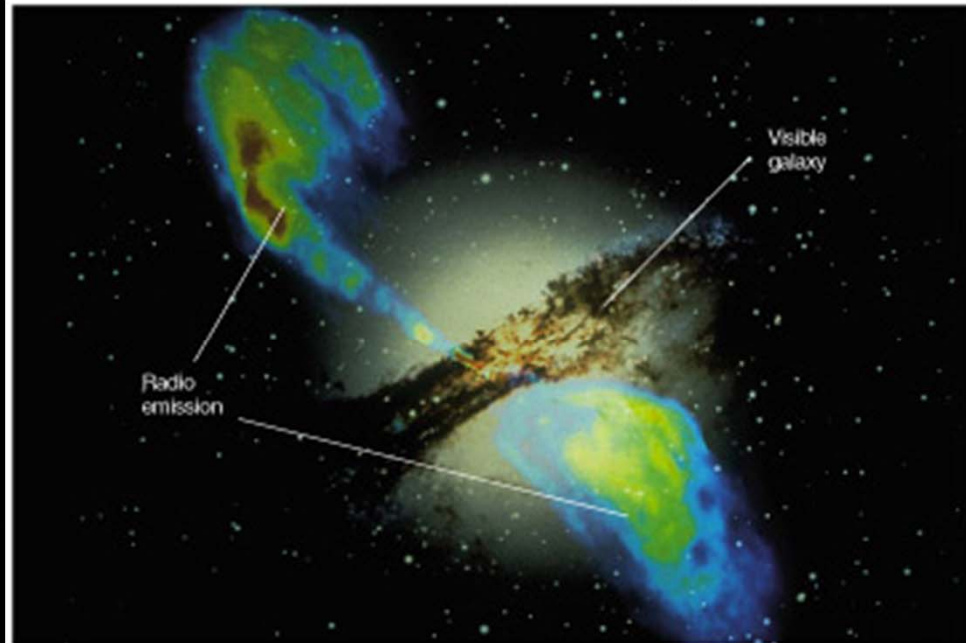
# RADIO TELESCOPES



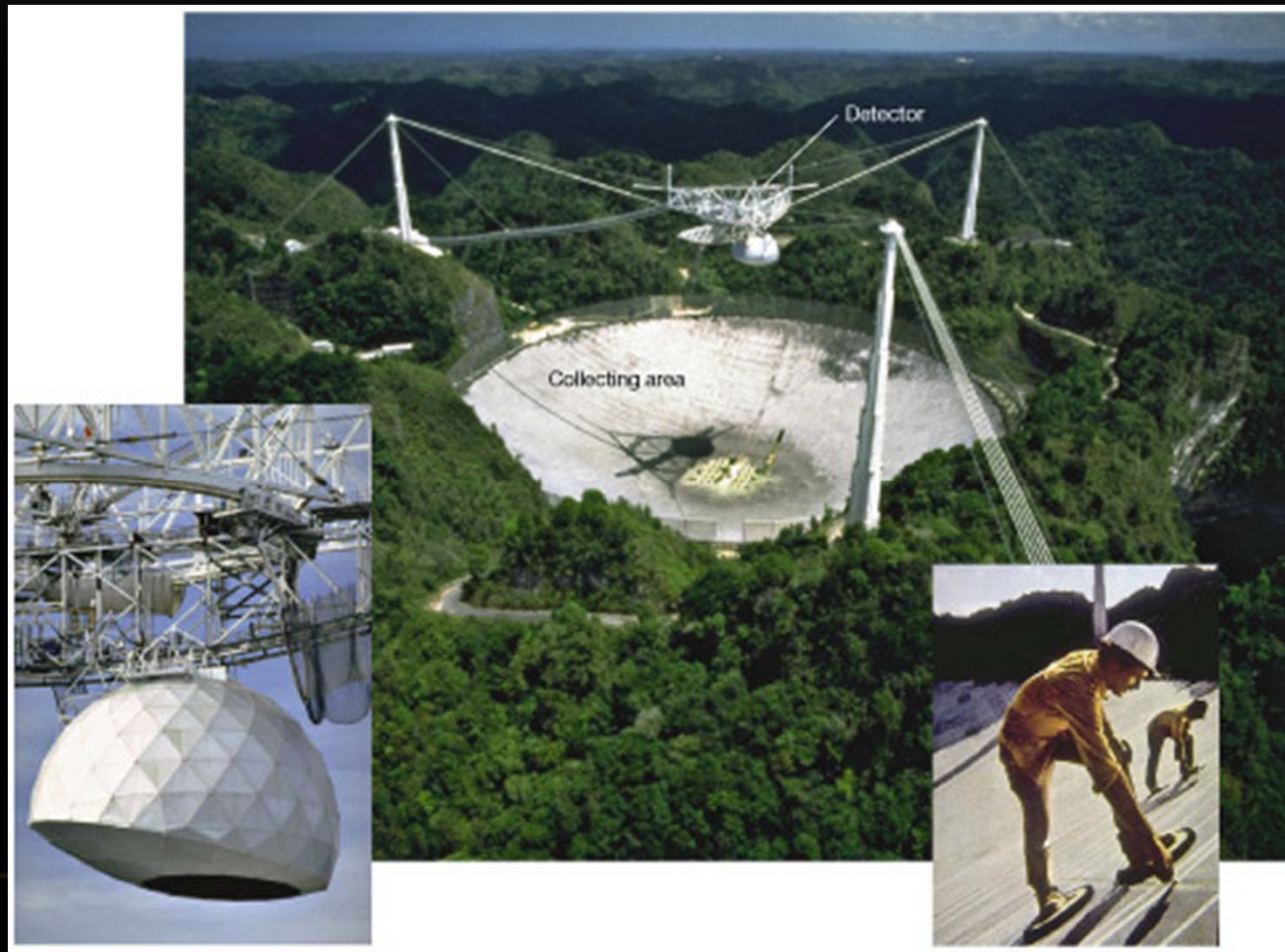
a)



(b)



# RADIO TELESCOPES





# INTERFEROMETRY

- Use an array of smaller scopes to get the angular resolution of a larger one



Fig.5.22

The Atacama Large Millimeter/submillimeter Array (ALMA) in Chile

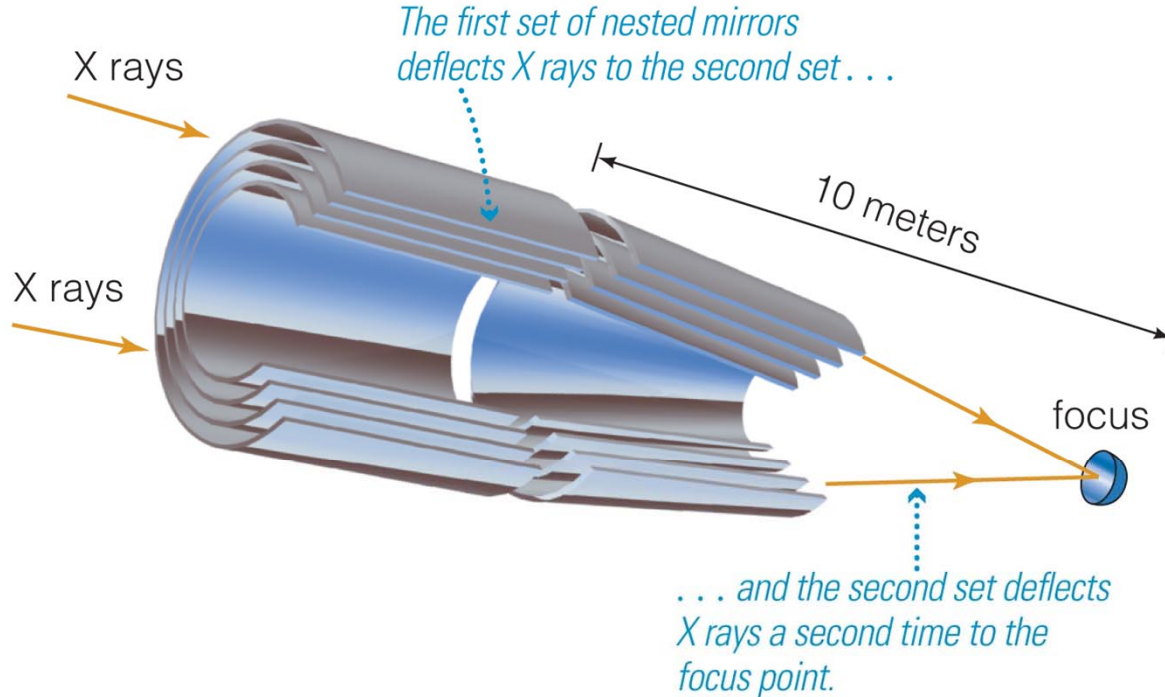


# X-RAY



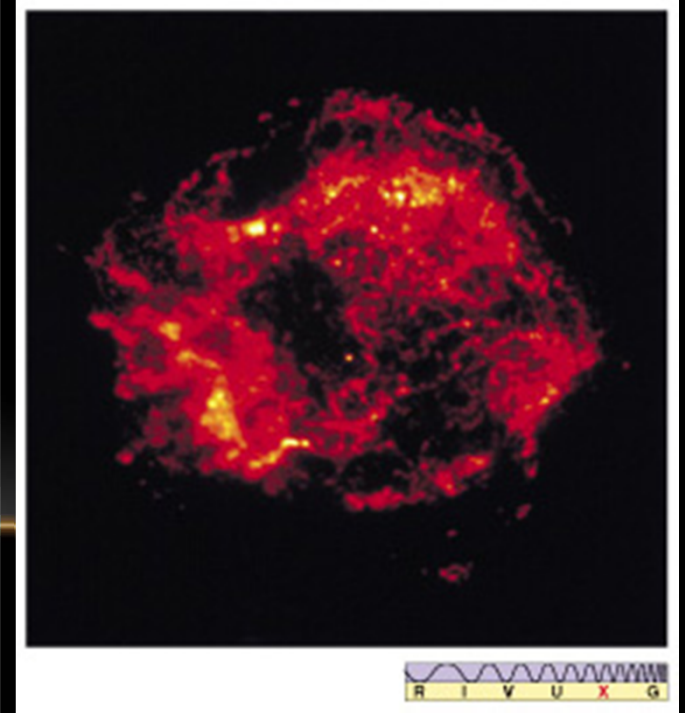
a Artist's illustration of the Chandra X-Ray Observatory, which orbits Earth.

Fig.5.23



b This diagram shows the arrangement of Chandra's nested, cylindrical X-ray mirrors. Each mirror is 0.8 meter long and between 0.6 and 1.2 meters in diameter.

A supernova remnant seen by Chandra in X-rays: click on [this link](#) for how it works

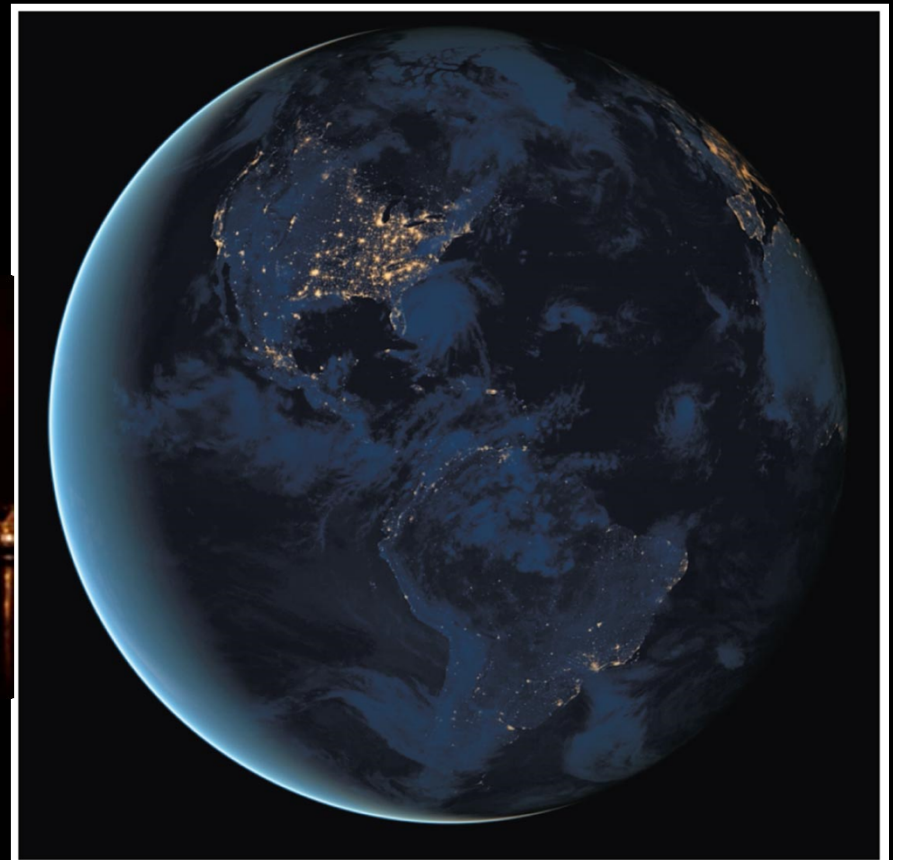


# WHERE TO PUT TELESCOPES?

- Problems caused by where you put a scope include light pollution
  - So location needs to be remote



Fig.5.24



# THE ATMOSPHERE...

- is not transparent to many wavelengths

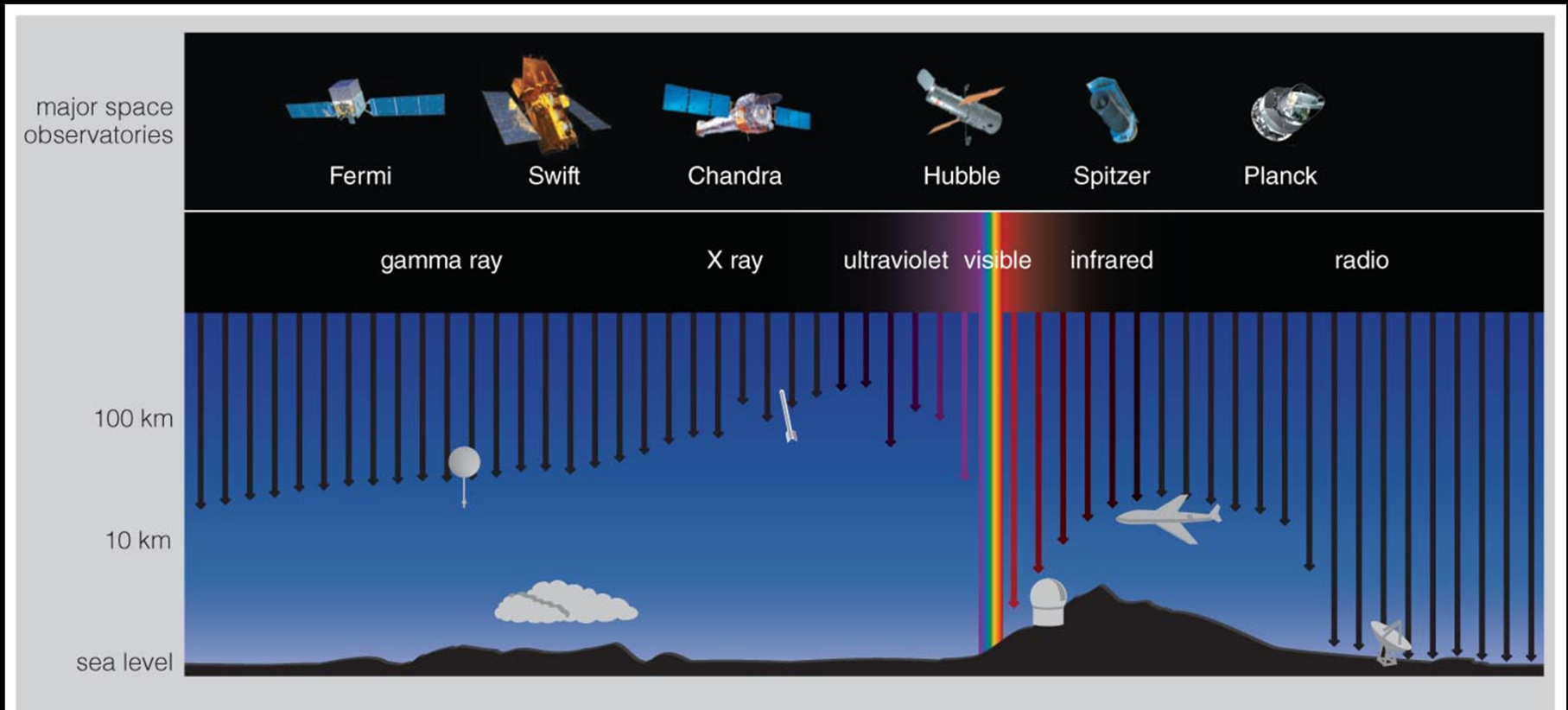
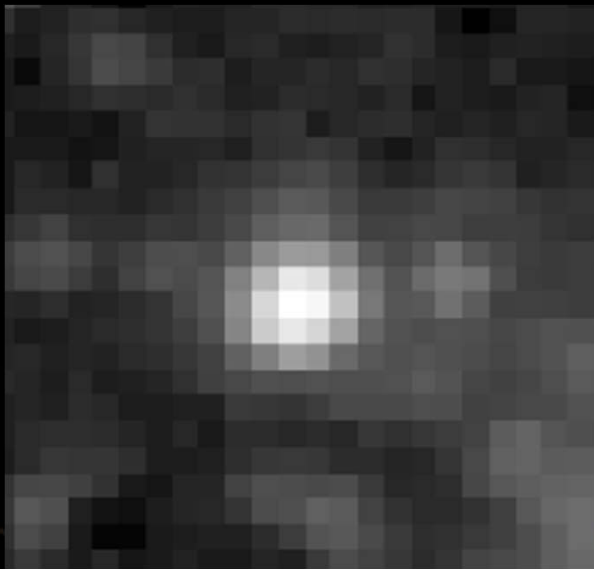


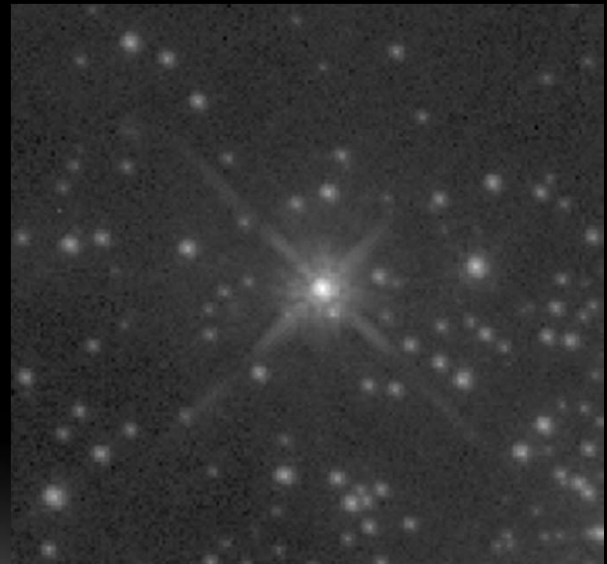
Fig.5.27

# EVEN IF TRANSPARENT...

- The air still smears out what you see
  - Same thing happening as when you see the wavyness in the air above a hot grill
  - Makes stars "twinkle"



Star viewed with ground-based telescope



View from Hubble Space Telescope



# SO, AVOID AS MUCH AIR AS POSSIBLE (*AND WEATHER!*)

- Go up a mountain to get above most of it
- Or in space to be above all of it



Fig.5.25  
Mauna Kea  
In Hawaii



Fig.5.26  
Hubble Space Telescope  
In low earth orbit